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ON THE WORLD DISTRIBUTION OF INCOME

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In this paper we demonstrate that the size distribution of the world income may be reasonably approximated by a log-normal distribution rather then by a power law, as has previously been believed. This result has been shown to be quite persistent as we move from 1985 to 2011.

JEL Codes: C13, C52, E25

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1. INTRODUCTION

With reference to the world income distribution, Di Guilmi *et al.* (2003) and Furceri (2008) have lent support to a power-law distribution linking the GDP per capita with the rank-size rule. In particular, in Di Guilmi *et al.* (2003), the GDP per capita of countries between the 30th and the 85th percentiles of the distribution is shown to follow a Pareto distribution over the period 1960–97. Furceri (2008), instead, provides empirical evidence of regularity in the GDP per capita data for 175 countries, from 1980 to 2004, collected by the International Monetary Fund (IMF) (World Economic Outlook (WEO), 2005). He shows that the "long-run" (average) world income distribution can be well approximated by Zipf's law.¹ Results in both papers are derived through the ordinary least squares (OLS) procedure.

As discussed in Clauset *et al.* (2009) and Urzúa (2011), the use of OLS for testing Zipf's law introduces systematic biases into the value of the exponent and returns unsatisfactory answers because it gives no indication of whether the data obey a power law at all. Indeed, several distributions, from the standard exponential to the log normal, may show linearity in their right tail on a Zipf plot, even if they are not power laws.

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¹The terms Pareto distribution and Žipf's law (the latter is also known as rank-size distribution or rank-size rule), after the two early researchers who championed their study, refer to cumulative distributions with a power-law form. Since power-law cumulative distributions imply a power-law form, Zipf's law and Pareto distribution are effectively synonymous with power-law distribution. Zipf's law and Pareto distribution are effectively synonymous with power-law distribution. Zipf's law and Pareto distribution are effectively synonymous with power-law distribution. Zipf's law and Pareto distribution are effectively synonymous with power-law distribution. Zipf's law and Pareto distribution are effectively synonymous with power-law distribution is plotted: Zipf made his plots with x on the horizontal axis and F(x) on the vertical one, Pareto did it the other way around (Newman, 2006, p. 4).

The issue is important because only a few distributions are consistent with Gibrat's law of Proportionate Effect (Gibrat, 1931), which provides the classical explanation for the empirical distribution of income. In particular, while the log-normal distribution can be derived as an outcome of stochastic growth models that rely on Gibrat's law (Sutton, 1997), the Pareto distribution is not consistent with it (Bottazzi, 2009).

This paper presents new evidence on the world income distribution. As in earlier studies, we show that the distribution of income is right skewed but, as a new result, we demonstrate that data distribute following a log-normal function rather than a power-law formulation. We demonstrate this point using a parametric method to estimate the income distribution for about 170 countries between 1985 and 2011 (IMF, WEO database, April 2013). For our analysis we use the three-parameter generalized gamma (GG) distribution, which includes the log-normal and the Pareto distributions as limiting cases (Kleiber and Kotz, 2003).

In the next section we present the database used for our analysis. The basic characteristics of the GG distribution and our estimation results are discussed in Section 3. Section 4 is devoted to assessment of the goodness of fit. Section 5 concludes.

2. DATA DESCRIPTION AND BASIC STATISTICS

In our study the world distribution of income is measured in terms of GDP per capita based on purchasing-power-parity (PPP). The unbalanced panel we use has been taken from the IMF, WEO database.² It contains selected macroeconomic data series from the statistical appendix of the WEO report, which presents the IMF staff's analysis and projections of economic developments at the global level in major country groups and in many individual countries.

Most countries' macroeconomic data presented in the WEO conform broadly to the 1993 version of the System of National Accounts (SNA)³ and reflect information from both national source agencies and international organizations. The WEO database is created during the biannual WEO exercise, which begins in January and June of each year, and results in a publication twice a year (April and September).

Our analysis refers to the April 2013 country-level database, which is made up of an average of 170 countries per year, over the time horizon 1985–2011. For every year in the sample, Table 1 shows the number of countries in the panel, some descriptive statistics, and the income concentration as measured by the Gini coefficient. Not surprisingly, the income data exhibit positive skewness. Moreover, for all the years in the panel, the empirical distribution of the world income looks leptocurtic. The Gini index shows a quite large income gap even though the inequality shows a slow tendency to decrease over the time horizon under consideration. In the same period, the median and mean have been increasing instead.

²http://www.imf.org/external/ns/cs.aspx?id=28 (accessed June 27, 2013).

³For details, see http://www.imf.org/external/pubs/ft/weo/2013/01/pdf/statapp.pdf.

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Year	No. countries	Min	Median	Mean	S.D.	Skew	Kur	Max	Gini		
1985	138	176.70	2,873	5,687	6,854.35	2.196	9.258	39,680	0.574		
1986	139	174.30	2,822	5,807	6,856.94	2.027	8.329	40,050	0.572		
1987	139	203.80	3,011	6,069	7,081.24	1.895	7.378	39,900	0.571		
1988	139	226.30	3,349	6,340	7,287.41	1.821	7.098	41,520	0.567		
1989	139	248.40	3,357	6,749	7,844.09	1.793	6.732	43,670	0.572		
1990	145	258.60	3,242	6,850	7,830.97	1.546	4.956	36,240	0.571		
1991	146	280.00	3,384	7,083	8,073.54	1.548	5.000	37,660	0.570		
1992	162	266.60	3,548	7,043	8,129.12	1.676	5.521	39,370	0.570		
1993	164	281.80	3,727	7,259	8,411.52	1.692	5.540	39,430	0.570		
1994	166	267.80	3,763	7,474	8,773.37	1.698	5.516	40,490	0.576		
1995	168	266.40	3,920	7,768	9,018.78	1.652	5.311	42,160	0.573		
1996	168	261.80	4,124	8,095	9,331.83	1.609	5.090	43,120	0.572		
1997	170	246.80	4,424	8,461	9,746.16	1.658	5.446	47,390	0.570		
1998	173	240.60	4,462	8,593	9,943.30	1.678	5.580	49,680	0.570		
1999	174	229.20	4,784	8,888	10,307.36	1.693	5.647	50,680	0.570		
2000	182	213.20	5,109	9,302	10,869.82	1.756	6.002	55,410	0.572		
2001	182	213.40	5,483	9,624	11,165.46	1.754	6.039	57,450	0.569		
2002	183	217.90	5,714	9,874	11,413.01	1.743	6.018	60,130	0.568		
2003	185	228.50	5,807	10,250	11,808.88	1.734	5.942	61,630	0.567		
2004	185	242.70	5,861	10,980	12,620.08	1.791	6.467	67,770	0.565		
2005	185	261.70	6,197	11,660	13,186.24	1.718	6.075	70,460	0.561		
2006	185	276.90	6,685	12,520	14,091.08	1.740	6.299	75,170	0.559		
2007	185	294.00	7,625	13,360	14,844.16	1.705	6.216	81,100	0.555		
2008	185	309.70	7,961	13,820	15,046.98	1.646	5.950	80,860	0.548		
2009	185	311.90	7,908	13,390	14,417.90	1.685	6.221	77,450	0.542		
2010	185	328.90	8,289	13,920	15,083.78	1.802	7.077	88,310	0.541		
2011	185	348.50	8,604	14,440	15,753.80	1.899	7.874	97,990	0.542		

TABLE 1 BASIC DESCRIPTIVE STATISTICS OF DATA

3. The GG and the World Distribution of Income

A positive random variable X is said to follow a $GG(a, \beta, p)$ if its probability density function (pdf) is given by

(1)
$$f(x) = \frac{ax^{ap-1}e^{-(x/\beta)^a}}{\beta^{ap}\Gamma(p)}, \quad x > 0.$$

Here $\Gamma(\cdot)$ is the standard gamma function, $\beta^a = \sigma^2 a^2$ is a scale parameter, whereas a > 0 and $p = \frac{a\mu + 1}{\beta^a}$ are shape parameters⁴ (McDonald, 1984; McDonald and Xu, 1995; Kleiber and Kotz, 2003). This parameterization was first introduced by Amoroso (1925) as the family of GG distributions.

The GG is a fairly flexible family of distributions that includes many distributions supported on the positive halfline as special or limiting cases (McDonald, 1984; McDonald and Xu, 1995; Kleiber and Kotz, 2003). In particular, as $a \to 0$, $p \to \infty$, $\beta \to 0$, but $a^2p \to 1/\sigma^2$ and $\beta p^{1/a} \to e^{\mu}$, the distribution tends to a log normal $LN(\mu, \sigma)^5$; for $a \to \infty$, $p = r/a \to 0$, with r > 0, the distribution tends to a power

⁴It is sometimes convenient to allow for a < 0. One then simply replaces *a* by |a| in the numerator of (1).

⁵Kleiber and Kotz (2003, p. 149) inadvertently indicate that $\beta \to \infty$, $a^2 \to 1/\sigma^2$ and $\beta p^{1/a} \to \mu$, as $a \to 0$.

function distribution PF $(r, \beta)^6$ (McDonald *et al.*, 2013). Since the power function distribution is the inverse Pareto distribution (Kleiber and Kotz, 2003, p. 73), for $a \to \infty$, $p \to 0$, with ap = -r, the Pareto distribution Par (β, r) is a limiting case as well. This makes the GG useful for discriminating among these models.

The GG was used by Kloek and Van Dijk (1978), Taillie (1981), McDonald (1984), Atoda *et al.* (1988), and Bordley *et al.* (1996) to study the income distribution in a number of different countries. In this paper, we use the GG to model the size distribution of the world income. The fitted model is given by $X \sim GG(\hat{a}, \hat{\beta}, \hat{p})$, where $\hat{a}, \hat{\beta}, \hat{p}$ are maximum likelihood estimates of *a*, β , and *p*. For the estimation of the three unknown parameters, we use a generalization of the Rigby and Stasinopoulos (1996a, 1996b) algorithm, which does not require accurate starting values to ensure convergence.⁷

Results of fitting the GG to our dataset are reported in Table 2. We give the estimates for the whole time window to analyze the evolution of the world income distribution over time. Values of \hat{a} , $\hat{\beta}$, and \hat{p} suggest that the log-normal distribution may be a good model for our data. Indeed, $\hat{\beta}$ and \hat{p} respectively tend to zero and infinity, namely to the limit value for the GG to converge to the log-normal function. \hat{a} is very close to zero, though it tends to increase over the time period under consideration, suggesting that the agreement with the log normal may deteriorate over time.

The relative fits of the log normal and GG with each other and with the data will be further explored in the next section. For the case under study, an increase in a confirms that the world income distribution has become less concentrated over time.

4. The Goodness of Fit

In order to get a formal test of the appropriateness of the log-normal hypothesis versus the GG model, we perform a likelihood ratio (LR) test, which provides the basis for comparing the goodness-of-fit (GoF) of nested models.

The LR test statistic is

(2)
$$LR = -2\ln\left(\frac{L(\Theta_r)}{L(\Theta_g)}\right) = 2[\ln L(\Theta_g) - \ln L(\Theta_r)],$$

where $L(\cdot)$ is the likelihood function, and Θ_g and Θ_r denote maximum likelihood estimators of the general and restricted model, respectively. The null hypothesis that the distribution with fewer parameters fits the data as well as the more general model is rejected if the value of the test is greater than a chi-squared percentile with *k* degrees of freedom, where the percentile corresponds to the chosen confidence level. In fact, the test statistic is asymptotically distributed as a χ^2 random variable with degrees of freedom equal to the difference in the number of free parameters between the two models. Based on the LR test results shown in Table 3, the hypothesis that the GG is observationally equivalent to the log-normal function

⁶Kleiber and Kotz (2003, p. 149) indvertently indicate that the PF is a limiting case of the GG as $a \rightarrow 0$; the correct limit is $a \rightarrow \infty$.

 $^{^{7}}$ R users may use the gamlss package.

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ESTIMATES FOR THE GG DISTRIBUTION	1998	$\begin{array}{c} 0.095 \\ 0.000 \\ 30,712 \end{array}$	2011	$\begin{array}{c} 0.425 \\ 0.039 \\ 17,593 \end{array}$			1998	0.17 (0.68)	2011	4.02 (0.045)				1998	0.37	$0.93 \\ 0.36$	2011	0.46	0.26																																		
	1997	$\begin{array}{c} 0.083\\ 0.000\\ 34,620 \end{array}$	2010	$\begin{array}{c} 0.440 \\ 0.050 \\ 16,580 \end{array}$																												1997	0.13 (0.72)	2010	4.29 (0.038)				1997	0.23	$0.92 \\ 0.33$	2010	0.43 0.90	0.23									
	1996	0.064 0.000 41,605	2009	$\begin{array}{c} 0.440 \\ 0.050 \\ 15,872 \end{array}$																					1996	0.06 (0.74)	2009	4.21 (0.040)				1996	0.26	0.90 0.30	2009	0.39 0.89	0.26																
	1995	$\begin{array}{c} 0.046 \\ 0.000 \\ 54,994 \end{array}$	2008	$\begin{array}{c} 0.427 \\ 0.044 \\ 15,860 \end{array}$		TABLE 3 Normal Model Against the GG	1995	0.03 (0.86)	2008	4.05 (0.044)				1995	0.32	0.90 0.29	2008	0.42 0.90	0.26																																		
	1994	$\begin{array}{c} 0.020 \\ 0.000 \\ 116,478 \end{array}$	2007	0.359 0.010 16,607			1994	0.01 (0.92)						1994	0.34	0.91	2007	0.53	0.29																																		
	1993	$\begin{array}{c} 0.058 \\ 0.000 \\ 41,871 \end{array}$	2006	0.311 0.002 16,926 1			TABLE 3 LR Statistics of the Log-Normal Model Against the GG	1993	0.05 (0.82)	5 2007) 2.94 3) (0.09)			t Model	1993	0.33	$0.92 \\ 0.32$	2006	0.51	0.34																																	
	1992	$\begin{array}{c} 0.055 \\ 0.000 \\ 42,840 \end{array}$	2005 2	0.276 0 0.000 0 17,134 16				TABLE 3 LR Statistics of the Log-Normal Model Ag	TABLE 3 5-Normal Model Ag	TABLE 3 g-Normal Model Ag	TABLE 3 g-Normal Model Ag	TABLE 3 g-Normal Model Ag	TABLE 3 g-Normal Model Ac	TABLE 3 g-Normal Model Ag	TABLE 3 G-Normal Model Ac	TABLE 3 6-Normal Model Ag	TABLE 3 g-Normal Model Ac	TABLE 3 g-Normal Model Ac	Model Ac	Model Ac	Model Ac	Model Ac	Model Ac	1992	0.04 (0.84)	2006	2.30 (0.13)		TABLE 4 P-VALUES FOR THE FIT TO THE LOG-NORMAL MODEL 1989 1990 1991 1992 1993 0.17 0.24 0.24 0.43 0.33 0.90 0.89 0.89 0.92 0.92	$0.92 \\ 0.33$	2005	0.53																					
	1991	0.058 0.000 39,100		ABLI	TABLE 3																	1991	0.04 (0.84)	2005	1.82 (0.18)		TABLE 4	it to the I	1991	0.24	0.89 0.24																						
	1990	$\begin{array}{c} 0.049 \\ 0,000 \\ 44,531 \end{array}$	2004	0.253 0.000 16,967		OF THE LO			1990	$\begin{array}{c} 0.01 \\ (0.92) \end{array}$	2004	$ \begin{array}{c} 1.59 \\ (0.20) \end{array} $			for the F	1990	0.24	0.26 0.26	2004	0.57																																	
	1989	0.084 0.000 25,714	2003	$\begin{array}{c} 0.189\\ 0.000\\ 20,160\end{array}$		LR Statistics			LR STATISTICS	1989	0.09 (0.76)	2003	0.86 (0.35)			P-VALUES	1989	0.17	0.90 0.28	2003	0.51	0.39																															
	1988	0.111 0.000 19,348	2002	$\begin{array}{c} 0.169 \\ 0.000 \\ 21,235 \end{array}$						LR	1988	0.16 (0.69)	2002	0.66 (0.42)				1988	0.13	0.89 0.25	2002	0.44 0.93	0.35																														
	1987	0.090 0.000 22,008	2001	$\begin{array}{c} 0.147 \\ 0.000 \\ 23,552 \end{array}$																																							1987	0.11 (0.74)	2001	0.49 (0.48)				1987	0.16	0.92 0.27	2001
	1986	0.086 0.000 22,138	2000	$\begin{array}{c} 0.135 \\ 0.000 \\ 24,063 \end{array}$				1986	0.11 (0.74)	2000	0.41 (0.52)	entheses.			1986	0.17	0.92 0.28	2000	0.51	0.40																																	
	1985	0.068 0.000 27,371	1999	$\begin{array}{c} 0.102 \\ 0.000 \\ 29,943 \end{array}$												1985	0.06 (0.81)	1999	0.21 (0.65)	p-Values in parentheses			1985	0.12	0.92	1999	0.44 0.93	0.40																									
	Year	$\hat{p}, \hat{\mathcal{B}}, \hat{a}$	Year	$\hat{p}, \hat{\mathcal{B}}, \hat{a}$			Year	LR	Year	LR	p-Va			Year	KS	AD CvM	Year	KS 4D	CvM																																		

TABLE 2

can not be rejected at conventional levels of significance, except for the end of the time period being considered, which is consistent with the estimated values of *a* increasing.

Finally, our log-normal hypothesis is investigated using three GoF tests: the Kolmogorov–Smirnov (KS), the Anderson–Darling (AD), and the Cramér–von Mises (CvM) tests.

For a hypothesized cumulative distribution function (CDF), F(x), the KS statistic is

$$D_n = \sup_{x} |F_n(x) - F(x)|,$$

where $F_n(x) = \frac{1}{n} \sum_{i=1}^{n} I_{X_i \le x}$ is the empirical cumulative distribution function (ECDF) ($I_{X_i \le x}$ is the indicator function that is equal to 1 if $X_i \le x$ and 0 otherwise). The null hypothesis that the sample comes from F(x) is rejected at level α if $\sqrt{n}D_n \ge K_{\alpha}$, where K_{α} is calculated by $P(K \le K_{\alpha}) = 1 - \alpha$, according to the Kolmogorov distribution (D'Agostino and Stephens, 1986).

Rather than computing the largest vertical distance between the two compared CDFs, the AD statistic is calculated over the whole range of data. This gives more weight to outliers than KS and makes the AD GoF test more powerful than KS for detecting departures in the tails from the hypothesized distribution.

The AD test statistics is

$$(4) AD = -n - S,$$

with

(5)
$$S = \sum_{i=1}^{n} \frac{2i-1}{n} [\ln(F(x_i)) - \ln(1 - F(x_{n-i+1}))].$$

Differently from KS, the AD test makes use of the specific distribution for calculating critical values. This is an advantage for sensitivity, but it requires that critical values be calculated for every single distribution tested. Several tabulated values can be found in D'Agostino and Stephens (1986).

An alternative to KS is the CvM statistic given by

(6)
$$CvM = \frac{1}{12n} + \sum_{i=1}^{n} \left[\frac{2i-1}{2n} - F(x_i)\right]^2.$$

Critical values for this test can be found in D'Agostino and Stephens (1986) and Anderson (1962). When KS and CvM tests give opposite results, the second one should be trusted (Anderson, 1962).

When, as in our study, parameters of the tested distribution are not known in advance but have to be estimated from the sample itself, simulation experiments or other methods are required to correct the test statistic and the critical values.

We have determined critical values following a parametric bootstrap procedure. We have first fit a log-normal model to our observed data to calculate parameters $\hat{\mu}$ and $\hat{\sigma}$, and then the test statistic for the observed data has been compared against the test value calculated for a large number (2500) of synthetic data sets, as large as the real data, drawn from a log-normal distribution with

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mean $\hat{\mu}$ and standard deviation $\hat{\sigma}$. The fraction of synthetic measurements larger than the empirical one defines our p-value.

Table 4 shows, for every year in the panel, the statistics for the fit to the log-normal model. p-Values are not statistically significant for every year under study and GoF test statistic, indicating that our dataset is consistent with a log-normal distribution.

5. CONCLUSIONS

In this paper we have analyzed the size distribution of the world income measured in terms of GDP per capita (PPP), using data taken from the IMF, WEO database (April 2013), for the period 1985–2011.

We have shown that the world income has a right-skewed distribution with an inequality Gini coefficient ranging between 0.541 and 0.576. For all the years in the panel we have also provided a fit to a three-parameter GG function, which includes the log-normal and the Pareto distributions as limiting cases.

The estimated values of parameters suggest that the world income distribution may be reasonably approximated by a log-normal model rather than by a power law. This result has been shown to be quite persistent as we move from 1985 to 2011. Goodness of fit tests support this finding.

The classic explanation for the empirical distribution of income is Gibrat's law of Proportionate Effect (Gibrat, 1931). In fact, the log-normal distribution can be derived as an outcome of stochastic growth models that rely on that famous empirical law (Sutton, 1997). This being the case, the possibility of Pareto's law shaping the size distribution of the world income has to be excluded because, as demonstrated in Bottazzi (2009), the law of Pareto and the law of Gibrat can not, in any respect, be reconciled.⁸

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⁸In this regard, conclusions of Di Guilmi et al. (2003) and Furceri (2008) should be reformulated.

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